

H160 Constants & Basic Formulae

k = kilo = 10^3	G = $6.67 \times 10^{-11} \text{ m}^3\text{kg}^{-1}\text{s}^{-2}$	h = $6.63 \times 10^{-34} \text{ Js}$	1 AU = $1.5 \times 10^7 \text{ km}$
M = mega = 10^6	m = milli = 10^{-3}	$\sigma = 5.65 \times 10^{-34} \text{ W/m}^2\text{/K}^4$	1 ly = $9.5 \times 10^{12} \text{ km}$
n = nano = 10^{-9}	c = $3 \times 10^8 \text{ m/s}$	k = $1.38 \times 10^{-23} \text{ J/K}$	1 pc = 3.26 ly

$\alpha = \frac{s}{d} 57.3^\circ$	$p^2 = a^3$	$v = \frac{\Delta x}{\Delta t}$	$a = \frac{\Delta v}{\Delta t}$	$net F = ma$	$p = mv$	$L = mvr$	$KE = \frac{1}{2}mv^2$
<i>small angle formulae</i>	<i>Kepler's 3rd law</i>	<i>velocity</i>	<i>acceleration</i>	<i>Newton's 2nd law</i>	<i>momentum</i>	<i>angular momentum</i>	<i>kinetic energy</i>

$E = mc^2$	$F = G \frac{M_1 M_2}{d^2}$	$a^3 = \frac{G}{4\pi^2} (M_1 + M_2) p^2$	$a^3 = (M_1 + M_2) p^2$	$v_{orbit} = \sqrt{\frac{GM}{R}}$	$v_{escape} = \sqrt{\frac{2GM}{R}}$
<i>mass energy</i>	<i>gravity</i>	<i>Newton's form Kepler's 3rd</i>	<i>Stellar Astronomer's form</i>	<i>orbital velocity</i>	<i>escape velocity</i>

$PV = NkT$	$D = \frac{m}{V}$	$c = \lambda f$	$E = hf = \frac{hc}{\lambda}$	$flux = \sigma T^4$	$\lambda_{max} = \frac{2.9 \times 10^6}{T} (nmK)$	$v_{rad} = \frac{\lambda_{shift} - \lambda_{rest}}{\lambda_{rest}} c$
<i>ideal gas law</i>	<i>density</i>	<i>wave equation</i>	<i>photon energy</i>	<i>radiation law #1</i>	<i>radiation law #1</i>	<i>Doppler shift</i>

$\frac{LGP_A}{LGP_B} = \frac{diam_A^2}{diam_B^2}$	$\alpha = 2.5 \times 10^5 \frac{\lambda}{D}$	$\gamma = \frac{1}{\sqrt{1-v^2/c^2}}$	$t_{move} = \frac{t_{rest}}{\gamma}$	$l_{move} = \frac{l_{rest}}{\gamma}$	$m_{move} = m_0 \gamma$	$\Delta x \Delta p \leq \frac{h}{4\pi}$
<i>light gathering power</i>	<i>resolving power</i>	<i>relativity gamma</i>	<i>relativity time</i>	<i>relativity length</i>	<i>relativity mass</i>	<i>uncertainty in x & p</i>

$\Delta E \Delta t \leq \frac{h}{4\pi}$	$\frac{\partial^2 \Psi}{\partial x^2} + \frac{8\pi^2 m}{h^2} (E - V) = 0$	$d(pc) = \frac{1}{p''}$	$b = \frac{L}{4\pi d^2}$	$L = 4\pi d^2 b$	$L = 4\pi R^2 \sigma T^4$
<i>uncertainty E and t</i>	<i>Schrodinger equation</i>	<i>distance parallax</i>	<i>apparent brightness</i>	<i>luminosity #1</i>	<i>luminosity #2</i>

$m - M = 5 \log \left(\frac{d}{10} \right)$	$d = 10^{\left(\frac{m - M + 5}{5} \right)}$	$(M_A + M_B) = \frac{a^3}{p^2}$
<i>distance #1</i>	<i>distance #2</i>	<i>star masses</i>